

LECTURE NOTE
ON
LAND SURVEY-II (TH.1)
6TH SEMESTER IN CIVIL ENGG.



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-: TACHEOMETRY :-

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- It is a branch of surveying in which horizontal and vertical distances are determined by taking angular observation with an instrument known as tachometer.
- This surveying is adopted in rough and difficult terrain where direct levelling and chaining are either not possible or very tedious.
- This surveying is very rapid and a reasonable contour map can be prepared for investigation works within a short time on the basis of such survey.

Objectives :- Tacheometry is the preparation of contour maps or plans requiring both the horizontal as well as vertical control. Also, on surveys of higher accuracy, it provides a check on distances measured with the tape.

USES :

1. Preparation of topographic map where both horizontal and vertical distances are required to be measured.
2. Survey work in difficult terrain where direct methods of measurements are inconvenient;
3. Reconnaissance survey for highways and railways, etc.
4. Establishment of secondary control points.

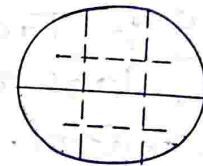
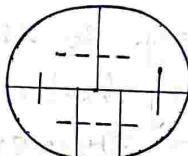
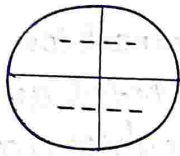
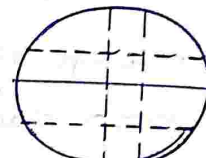
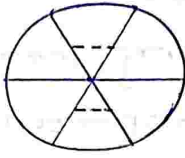
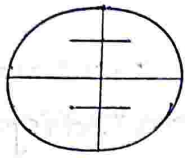
Tacheometric Surveying Instruments :-

Tacheometric surveying is done with the help of Tachometer and Stadia Rod.

1. Tachometer
2. Stadia Rod
3. Anallatic lens.

1. Tachometer :- A transit theodolite fixed with special stadia diaphragm is known as tachometer. It is the main instrument of tacheometric surveying. Its telescope contains two horizontal hairs called stadia hairs in addition to the regular crosshair.

The stadia hairs are equidistance from the central cross-hairs and they are specially termed as stadia lines or stadia webs. The common types of stadia diaphragms are shown below:



2. Stadia Rod :- For small distances (up to 100 meters) a level staff may be used for tachometric surveying. But for greater distances stadia rod is needed. Stadia rod is of one piece having 3 to 5 meter length. The smallest subdivision is usually 5 mm.

3. Anallatic lens :- It is an additional lens used in the instrument. It is a special lens which is placed between the object glass and the eyepiece of the telescope in order to eliminate the additive constant ($f + d$). This is done to make the expression for the distance between instrument station and staff position more simplified. The lens is only provided in an external focusing telescope but not in the internal focusing.

Features of Tacheometer :-

1. Multiplying constant value of 100 and additive constant zero.
2. Axial horizontal line should be exactly midway between the other two lines.
3. Telescope should be truly anallatic.
4. Telescope should be powerful, having a magnification of 20 to 30 diameters.

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5. Aperture of the objective should be 35 to 45mm in diameter to have a bright image.
6. For small distances (upto 100 meters), ordinary levelling staff may be used, for greater distances a stadia rod may be used.
7. A stadia rod is usually of one piece, having 3 to 5 metre length.
8. A stadia rod graduated in 5mm (i.e., 0.005m) for smaller distances and while for longer distances, the rod may be graduated in 1cm (i.e., 0.01m).

Principle of stadia method :-

The stadia method is based on the principle that the ratio of the perpendicular to the base is constant in similar isosceles triangles. Two rays OA and OB be equally inclined to central ray OC. Let A_2B_2 , A_1B_1 and AB be staff intercepts.

from the figure we can write,

$$\frac{OC_2}{A_2B_2} = \frac{OC_1}{A_1B_1} = \frac{OC}{AB} = \frac{f}{i} = K$$

where $\frac{f}{i} = K$ is multiplying constant.

f = focal length of the objectives

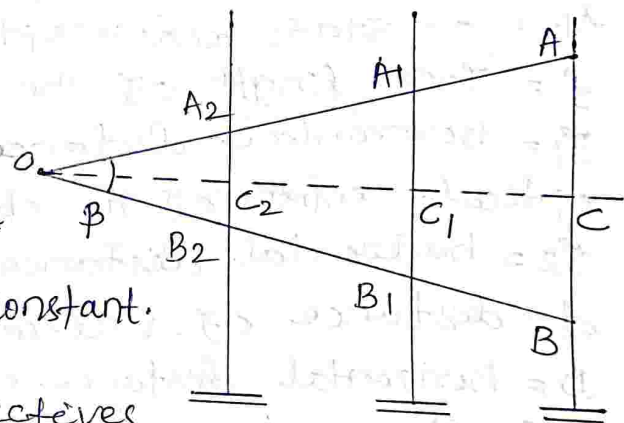
i = interval between the stadia hairs (stadia interval)

This constant K entirely depends upon the magnitude of the angle B .

→ In actual practice, observations may be made with either horizontal line of sight or with inclined line of sight.

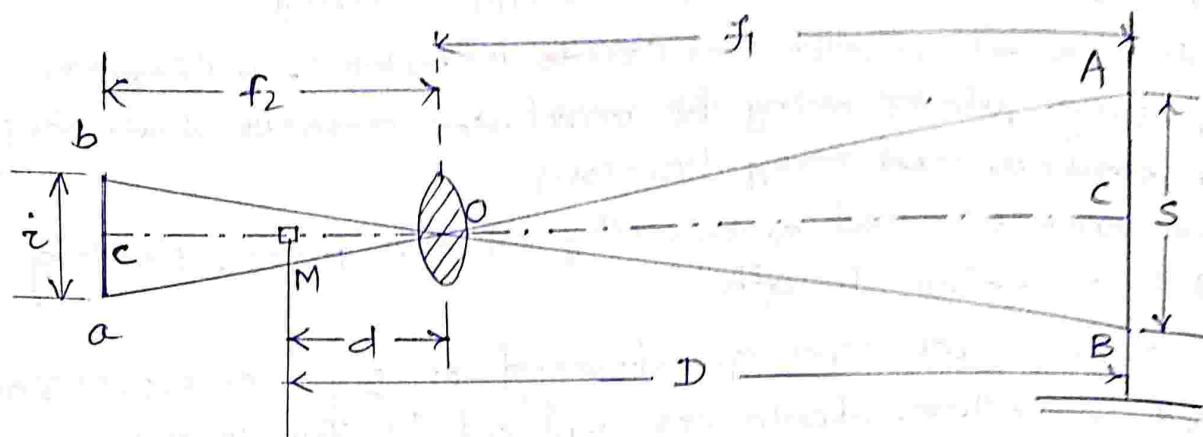
→ In latter case the staff may be kept either vertically or normal to the line of sight.

→ First the distance-elevation formula for the horizontal sights should be derived.



Horizontal sights:

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Consider the figure, in which O is the optical centre of the objective of an external focusing telescope.

Let A, C and B are the points cut by the three lines of sight corresponding to three wires. b, c and a are the top, axial and bottom hairs of diaphragm.

$ab = i =$ interval between the stadia hairs.

$AB = s =$ staff intercept (stadia interval)

$f =$ Focal length of the objective.

$f_1 =$ horizontal distance of the staff from the optical centre of the objective

$f_2 =$ horizontal distance of the cross-wires from O.

$d =$ distance of vertical axis of instrument from O

$D =$ horizontal distance of staff from vertical axis of the instrument.

M = Centre of the instrument, corresponding to the vertical axis.

Since the rays BOB and AOa pass through the optical centre, they are straight, so that AOB and aOb are similar, hence,

$$\frac{f_1}{f_2} = \frac{s}{i}$$

Again since, f_1 and f_2 are conjugate focal distance we have from lens formula,

$$\frac{1}{f} = \frac{1}{f_2} = \frac{1}{f_1}$$

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Multiplying throughout by $\frac{f}{f_1}$, we get $f = \frac{f_1}{f_2} f + f$
 Substituting the values of $\frac{f_1}{f_2} = \frac{S}{2}$ the above equation
 we get, $f = \frac{S}{2} f + f$

Horizontal distance between the axis and the
 staff is $D = f + d$

$D = \frac{f}{2} S + (f + d) = K \cdot S + C$, this equation
 is known as the distance equation.

The constant $K = \frac{f}{2}$ is known as the multiplying
 constant or stadia interval factor and the
 constant $(f + d) = C$ is known as the additive
 constant of the instrument.

Determination of constant K and C :

1st Method :- This is the laboratory method. In this
 method, the additive constant $C = (f + d)$ is measured
 from the instrument while the multiplying constant
 K is computed from field observation.

1. Focus the instrument to a distant object and
 measure along the telescope the distance between
 the objective and cross-hairs.

$$\frac{1}{f} = \frac{1}{f_1} = \frac{1}{f_2}$$

2. The distance d between the instrument axis and
 the objective is variable in the case of external
 focusing telescope, being greater for short sights
 and smaller for long sights. It should, therefore
 be measured for average sight. Thus the additive
 constant $(f + d)$ is known.

3. To calculate the multiplying constant K, measure
 a known distance D_1 and take the intercept
 S_1 on the staff kept at that point, the line
 of sight being horizontal. Using the equation,

$$D_1 = K S_1 + C \quad \text{or} \quad K = \frac{D_1 - C}{S_1}$$

For average values, staff intercepts, S_2, S_3 , etc can be measured corresponding to distance D_2, D_3 , etc. an mean value can be calculated.

2nd Method :- In this method, both the constants are determined by field observations as under:-

1. Measure a line about 200m long, on fairly leveled ground and drive pegs at some interval say 50 meters.
2. Keep the staff on the pegs and observe the corresponding staff intercepts with horizontal sight.
3. Knowing the values of D and S for different points, a number of simultaneous equations can be substituting the values of D and S in equation $D = ks + C$. The simultaneous solution of successive pairs will give the values of k and C , and the average of these can be found.

For example, if S_1 is the staff intercept corresponding to distance D_1 and S_2 corresponding to D_2 we have,

$$D_1 = kS_1 + C \text{ --- (I) and } D_2 = kS_2 + C \text{ --- (II)}$$

Subtracting (I) from (II) we get,

$$k = \frac{D_2 - D_1}{S_2 - S_1} \text{ --- (1)}$$

Substituting the value of k in (I), we get,

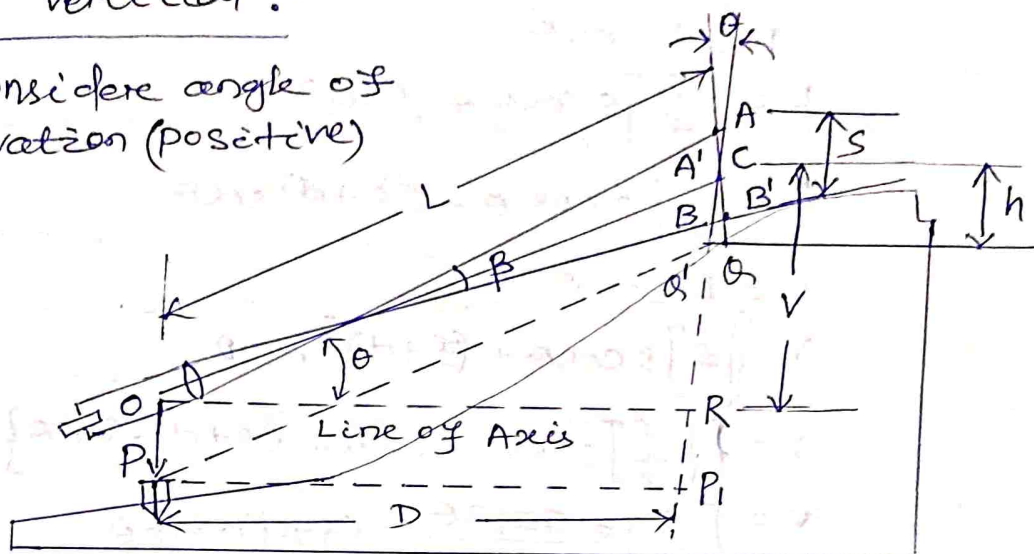
$$C = D_1 - \frac{D_2 - D_1}{S_2 - S_1} S_1 = \frac{D_1 S_2 - D_2 S_1}{S_2 - S_1} \text{ --- (2)}$$

Thus equation (1) and (2) give the values of k and C .

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Expression of Horizontal distance and vertical Elevation, if line of sight is inclined and staff is held vertical :-

Case (i) Consider angle of elevation (positive)



(Line of sight is inclined and staff held vertical)

Let P = Instrument station; Q = Staff station.

M = Position of instruments axis;

O = Optical centre of the three hairs.

$S = AB$ = Staff intercept; i = Stadia interval

θ = Inclination of line of sight from the horizontal

L = Length OC measured along the line of sight.

$D = OR$ = Horizontal distance between the instrument and staff.

V = Vertical intercept at Q , between the line of sight and the horizontal line.

h = height of the instrument;

r = central hair reading

β = Angle between two extreme rays corresponding to stadia hairs.

Draw a line $A'B'$ normal to line of sight OC . Angle

$\angle AA'C = 90^\circ + \beta/2$, being the exterior angle of the $\triangle COA'$

Similarly, from $\triangle COB'$, angle $\angle OB'C = \angle BB'C = 90^\circ - \beta/2$

Since $\beta/2$ is very small Cts value = $17''$ for $k=100$)

$\angle AA'C$ and $\angle BB'C = 90^\circ$

from $\triangle ACA'$, $A'C = AC \cos \theta$ or $A'B' = AB \cos \theta = S \cos \theta$

Since line $A'B'$ is perpendicular to line of sight OC , equation $D = ks + c$ is directly applicable.

We have, $S' = AC \cos \theta + BC \cos \theta$

$$S' = (AC + BC) \cos \theta$$

$$S' = S \cos \theta$$

$$D = L \cos \theta$$

$$L = \left[\frac{f}{2} \right] S \cos \theta + (f+d)$$

$$D = \left[\frac{f}{2} \right] S \cos^2 \theta + (f+d) \cos \theta$$

$$V = L \sin \theta$$

$$V = \left\{ \left[\frac{f}{2} \right] S \cos \theta + (f+d) \right\} \sin \theta$$

$$V = \left\{ \left[\frac{f}{2} \right] S \cos \theta \cdot \sin \theta + (f+d) \cdot \sin \theta \right\}$$

$$V = \left[\frac{f}{2} \right] S \frac{\sin 2\theta}{2} + (f+d) \cdot \sin \theta$$

* If line of sight is inclined upwards = H.I. of instrument axis + V - h

* If line of sight is inclined downwards = H.I. of instrument axis - V - h.

Case-iii) Consider angle of depression (negative).

$$L = \left(\frac{f}{2} \right) S \cos \theta + (f+d)$$

$$D = L \cos \theta$$

$$D = \frac{f}{2} S \cos^2 \theta + (f+d) \cos \theta$$

$$V = L \sin \theta$$

$$V = \left(\frac{f}{2} \right) S \cos \theta \cdot \sin \theta + (f+d) \sin \theta$$

$$V = \left(\frac{f}{2} \right) S \frac{\sin 2\theta}{2} + (f+d) \sin \theta$$

R.L of staff station = R.L of line of collimation - V - h

